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INTRODUCTION TO NUCLEAR DUST/DEBRIS CLOUD FORMATION. (U)

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# INTRODUCTION TO NUCLEAR DUST/DEBRIS CLOUD FORMATION

Martin Rosenblatt  
California Research and Technology, Inc.  
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1 July 1981

Topical Report for Period 7 November 1979—1 July 1981

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20. ABSTRACT (Continue on reverse side if necessary and identify by block number)  The detonation of a nuclear device on or near the Earth's surface generates a lofted cloud composed of dust and other debris. The objectives of this report are to provide (1) an overview of dust/debris cloud phenomenology and (2) examples of cloud characteristics. Physical phenomenology associated with surface bursts (e.g., crater ejecta) and airbursts (e.g., shock reflections and sweep-up layer formation) are discussed and examples are presented.		

## PREFACE

Questions concerning dust/debris environments generated by surface and near-surface nuclear detonations continue to be asked by defense planners. This report is intended to introduce the subject to those individuals not familiar with nuclear dust/debris cloud formation.

Lt. Col. R. E. Case (DNA/SPAS) was the technical monitor for this work and his helpful suggestions are gratefully acknowledged.

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**Table of Conversion Factors  
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Physical Quantity	To convert from units in report	To metric (SI) units	Multiply by
Mass	pound ton kton	kilogram (kg) kilogram (kg) kilogram (kg)	.454 $10^3$ $10^6$
Energy	ton kton Mton	joules (J) joules (J) joules (J)	$4.2 \times 10^9$ $4.2 \times 10^{12}$ $4.2 \times 10^{15}$
Pressure	bar* psi*	pascal (Pa) pascal (Pa)	$10^5$ 6895.
Density	gm/cm <sup>3</sup>	kg/m <sup>3</sup>	$10^3$
Length	foot	meter (m)	.3048

\*1 bar = 14.5 psi

1 psi = .06895 bar

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SECTION 1  
INTRODUCTION

1.1 BACKGROUND AND OBJECTIVES

The detonation of one or more nuclear weapons on or near the Earth's surface generate a lofted cloud composed of dust and other debris.

The lofted mass is a potential threat to

- *reentry vehicles* which must fly through clouds from prior detonations (fratricide)
- *ascent vehicles* which must flyout soon after being attacked
- *radar and communications* which must operate in a dust/debris environment

The objectives of this report are to provide interested persons with

1. an overview of dust/debris cloud phenomenology
2. examples of cloud characteristics

## 1.2 BASIC CONCEPTS

After a near-surface nuclear burst, the cloud characteristics of primary interest are the concentration\* and particle size distribution of the condensed phase mass in the cloud as a function of position and time.

The condensed phase mass in the cloud can be composed of:

- *earth material* which has been entrained and lofted by the ascending winds which develop. (Earth or ground material include soil/rock, water, vegetation, as well as structures near the burst point.)
- *weapon material* which has been vaporized and recondenses during the early phases of the nuclear cloud development.
- *water/ice particles* which condense from relatively moist (humid) air which is entrained and lofted from near the Earth's surface to higher and colder regions of the atmosphere.

Particle is a generic term which includes dust, pebbles, rocks, boulders, water, ice, or any other condensed phase mass in the cloud. The lofted particles are of many different sizes and shapes. For convenience, the size of a particle is usually taken as the equivalent spherical diameter,  $a$ :

$$a = \left( \frac{6m}{\rho_o \pi} \right)^{1/3} \quad \text{Diameter (m = 1 gm when } a = 1 \text{ cm and } \rho_o = 1.9 \text{ gm/cm}^3)$$

where

$m$  Mass of particle

$\rho_o$  Density of particle

---

\*Mass concentration is also commonly called density.

The weapon yield ( $W$ ) is the energy released in a nuclear detonation.\* The yield is usually expressed in equivalent tons of TNT, where

1 ton of TNT is assumed to release  $10^9$  calories or  
 $4.2 \times 10^9$  Joules

1 kiloton = 1 kt =  $4.2 \times 10^{12}$  Joules

1 megaton = 1 Mt =  $4.2 \times 10^{15}$  Joules

The first nuclear detonation, shot TRINITY in 1945, had a yield of 19 kt. The yields of the weapons detonated over Hiroshima and Nagasaki were approximately 13 kt and 23 kt, respectively. Shot CASTLE BRAVO in 1954 at Bikini (in the Pacific Proving Ground) had the largest U.S. test yield of ~15 Mt.

The height of burst (HOB) is the device height above the Earth's surface. The scaled height of burst (SHOB) is defined as

$$SHOB = HOB/W^{1/3} \quad (2-1)$$

Buried bursts have negative HOB and SHOB.

A contact surface burst is defined as  $|SHOB| < 5 \text{ ft}/\text{kt}^{1/3}$ . A nuclear detonation this close to the ground surface causes a dynamic cratering process which ejects large amounts of earth material into the atmosphere. A contact surface burst couples about 8% of the nuclear yield to the ground in a few microseconds. The details of the energy coupled to the ground, and the amount of ejecta thrown into the atmosphere will depend on  $W$ , HOB, the device construction, and the properties of the ground material.

---

\*The nature of the nuclear energy released and a general description of nuclear weapons effects are in Reference 1.

As the SHOB increases above  $5 \text{ ft}/\text{kt}^{1/3}$  (i.e., about 50 ft for  $W = 1$  megaton), the fraction of energy coupled directly to the ground and the amount of ejected mass begins to decrease rapidly. An estimate of the crater ejecta mass source vs HOB and SHOB for  $W = 1$  megaton is shown in Figure 1 below. All other sources of ground material which become entrained in the nuclear cloud come from near the ground surface and can be conveniently combined into the sweep-up mass source. The sweep-up mass source dominates the lofted cloud mass characteristics for  $\text{SHOB} \gtrsim 20 \text{ ft}/\text{kt}^{1/3}$ .

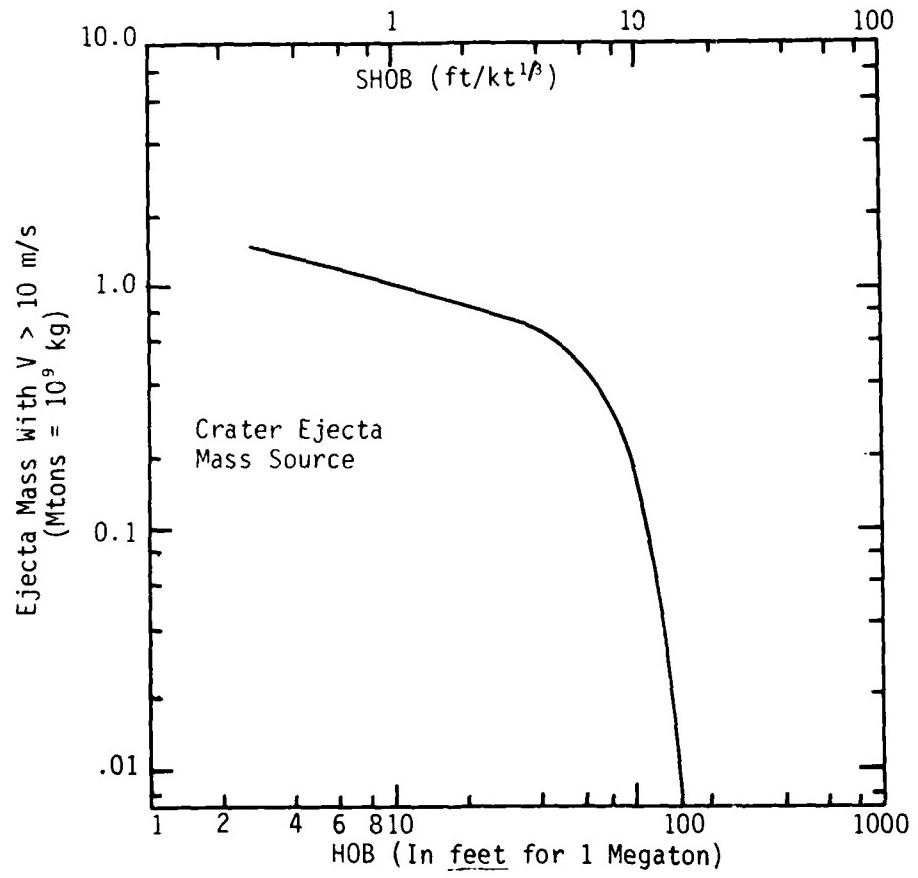


Fig. 1. Crater Ejecta Mass Source with Ejection Speeds of Over 10 m/s versus HOB for  $W = 1 \text{ Mt}$ .<sup>2</sup>

## SECTION 2

### DUST/DEBRIS SOURCES

#### 2.1 CRATER EJECTA MASS SOURCE

For contact surface bursts, crater ejecta represents the primary source of ground material which becomes entrained in the nuclear cloud. The following figure indicates the general physical phenomena, pressures, and times involved in crater formation for a nominal 1 megaton contact surface detonation.

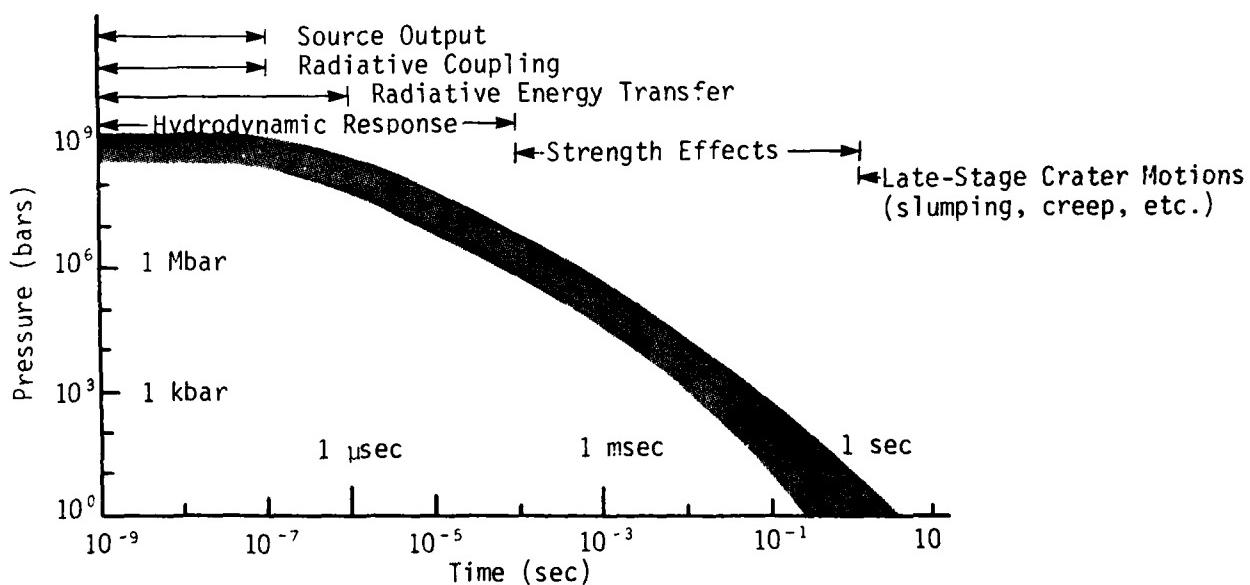


Fig. 2. Typical Pressure-Time Sequence for Energy Coupling and Cratering.

The following photos show the fireball and ejecta from the slightly buried burst JOHNIE BOY ( $W = 0.5$  kt, HOB  $\sim 2.0$  ft, at NTS). Note that some of the crater ejecta emerges from the early-time fireball in distinct rays.

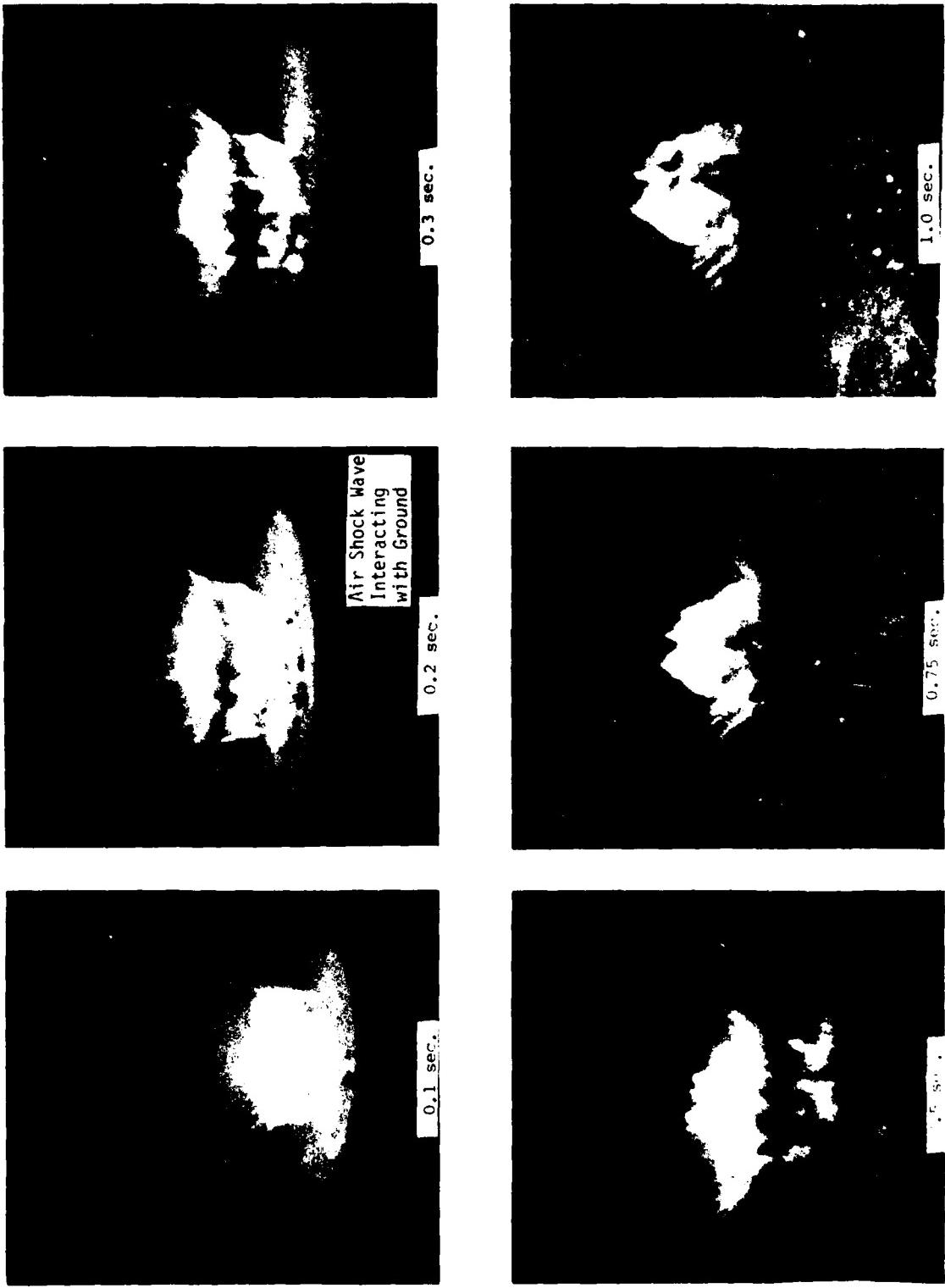
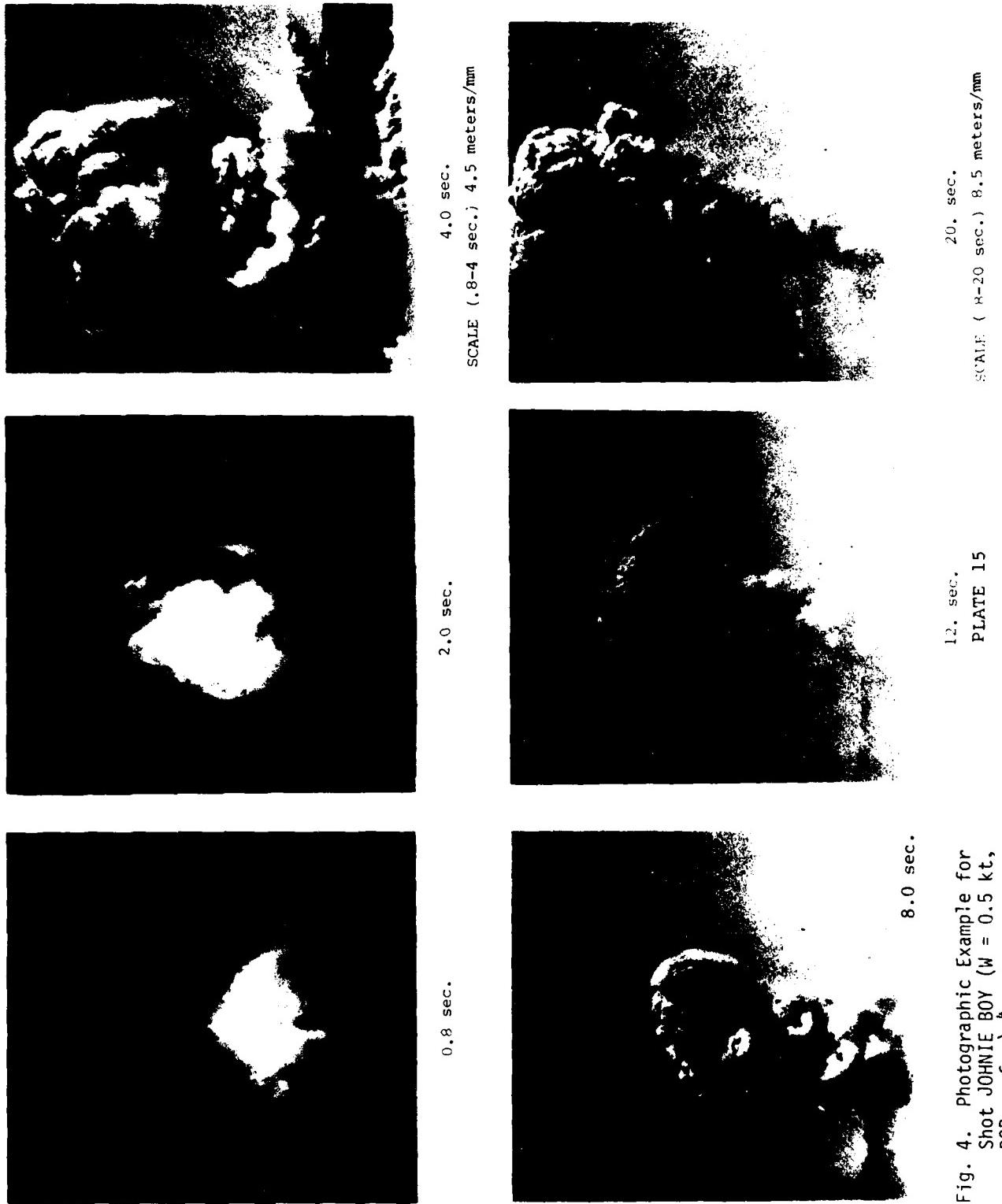


Fig. 3. Photographic Example for  
Shot JOHNNIE BOY ( $W = 0.5$  kt,  
 $DOB = .6$  m).<sup>4</sup>

PLATE 14

HORIZONTAL SCALE 5.7 meters/mm



Theoretical calculations (or computer code simulations) of the crater and ejecta formation use finite difference analogues of the differential equations for conservation of mass, conservation of momenta, conservation of energy, and approximate equations of state for the weapon and ground materials.

These numerical simulations predict the particle velocity, density and temperature of the incipient ejecta versus radius and time. (Most crater/ejecta calculations are axisymmetric and ejecta rays are ignored.) Figure 5 (next page) shows a velocity vector field near the dynamic crater at  $t \sim 0.2$  seconds for a 1 Mt contact burst over the indicated geology. The incipient ejecta is flowing through the original ground surface ( $Z = 0$ ) at radii between 40 m and 80 m at 0.2 seconds. The ejection angles are roughly  $45^\circ$ .

The ejecta mass rate per unit area flowing across  $Z = 0$  at any radius is  $\rho V$ . At  $t = 0.2$  seconds,  $\rho V$  peaks at a radius of about 50 meters and has an associated vertical velocity of  $\sim 60$  m/s.

Figure 6 shows the analogous vertical velocity (i.e., associated with the peak mass flux) as a function of time from 5 nuclear cratering calculations. The various calculations show variations of about a factor of 2 from the mean curve.

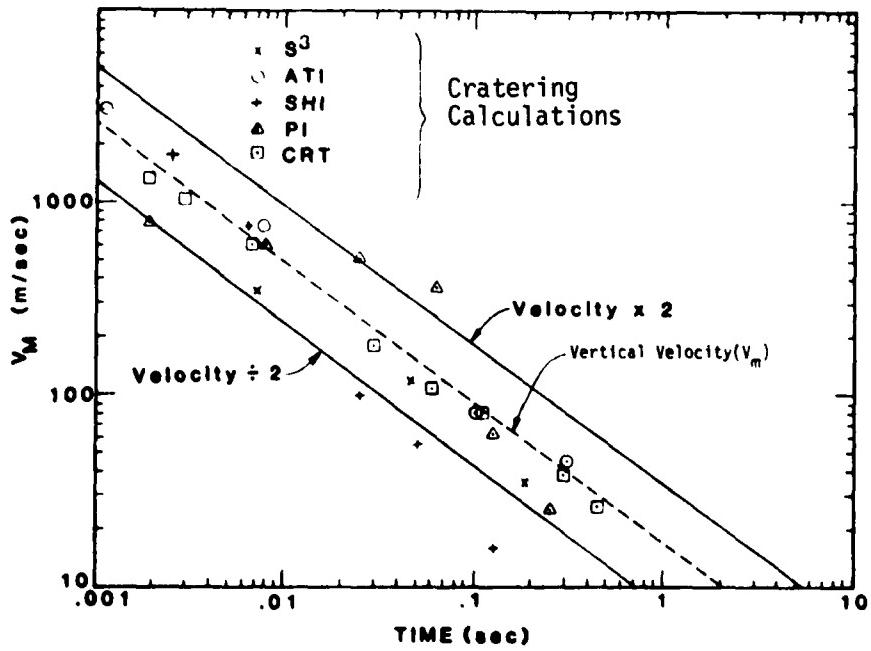


Fig. 6. Ejecta Vertical Velocity at a Range of Greatest Mass Flux from Cratering Calculations for a 1 Mt Surface Burst.<sup>5,6</sup>

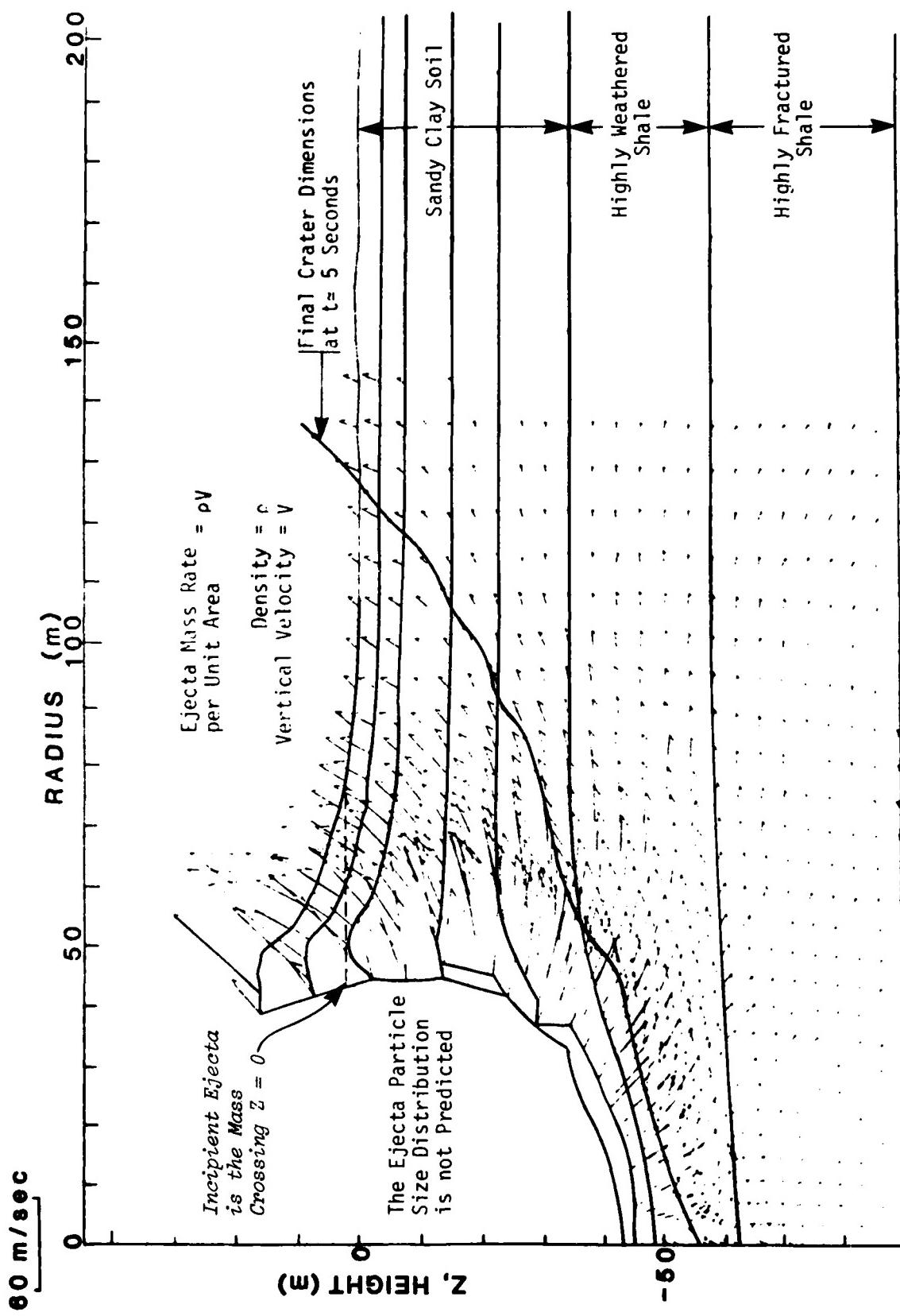


Fig. 5. Crater Ejecta Source Example: Velocity Field at  $t \sim 0.2$  Seconds for  
a  $W = 1$  Mt Surface Burst.<sup>6</sup>

Of particular importance (for the lofted mass characteristics of the nuclear cloud) is the *particle size distribution* of the incipient ejecta mass. This size distribution is *not* provided by current cratering calculations. Physically, the particle size distribution will depend on, at least, the stress, strain, and temperature history of each element of the incipient ejecta. In general, as the ejecta flow into the air, the particle size distribution will vary with both the time of ejection and the range from the burst point. The incipient ejecta initially consist of vaporous soil and bomb debris material. At somewhat later times, melted soil flows into the fireball. The bulk of the ejecta, however, is composed of solid material which has been shocked to some peak pressure, distorted, and flowed by the cratering dynamics. For example, about 500 kton of soil experiences 7 kbar (100 ksi) or more of pressure in a 1 Mt yield surface burst. Thus, the crater ejecta particle size uncertainty involves not only natural in-situ size variations, but also uncertainties in soil/rock break-up and agglomeration during crater formation.

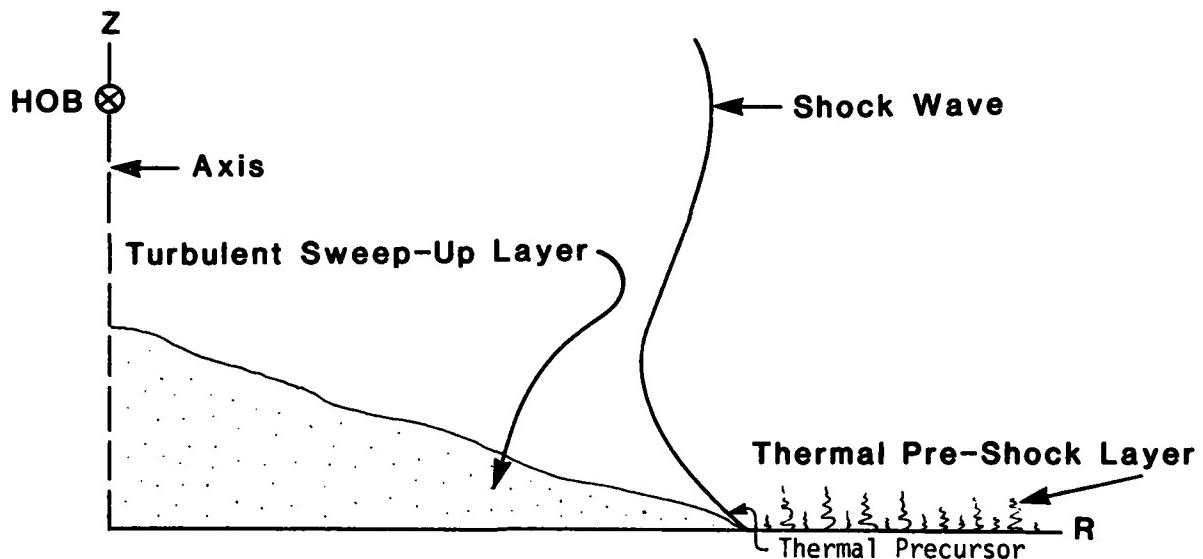
The crater does not completely form until about 5 seconds for a 1 Mt surface burst (an example of a predicted final crater profile is shown with a dashed line on Figure 5). Roughly 2 Mtons of ejecta mass ( $1 \text{ Mton} = 10^{12} \text{ gm}$ ) is thrown from the crater, however, note that only a small fraction\* of the ejecta mass becomes lofted as part of the nuclear cloud. Most of the mass falls back to earth between 10 seconds and 1 minute.

---

\*It is estimated that ~0.3 Mtons of ejecta mass per Mt of weapon yield remains aloft at 5 minutes after detonation of a contact surface burst.

## 2.2 SWEEP-UP MASS SOURCE

For non-cratering HOB\* detonations, the dust/pebbles which become entrained in the nuclear cloud originate very near the ground surface (probably within a few centimeters or so). A "dusty" sweep-up layer is formed as this near-surface material mixes with the high velocity winds behind the air shock wave. The sweep-up layer is fairly persistent, and some of the dust/pebble particles in this layer are subsequently drawn up into the main cloud.



Thermal heating of an air layer near the ground surface may occur prior to shock arrival; this heated layer causes a thermal precursor flow field which appears to enhance the sweep-up layer. A thermal precursor is a pressure and velocity wave disturbance which travels faster in the heated near-surface thermal layer than in the cooler air above the surface.

---

\*HOB is commonly used as an abbreviation for "height of burst" as well as the burst distance above the ground surface.

Figure 7 is a photographic example of a relatively late-time (~4 sec) sweep-up layer for shot GRABLE ( $W = 15$  kt, HOB = 524 ft, SHOB =  $212 \text{ ft}/\text{kt}^{1/3}$ ). Based on existing Nevada Test Site (NTS) data for SHOB between 150-500  $\text{ft}/\text{kt}^{1/3}$ , the maximum radial extent of the sweep-up layer is about  $R_{\max} \sim 1000-1300 \text{ ft}/\text{kt}^{1/3}$  which corresponds to a peak overpressure of ~5-10 psi. The maximum height of the sweep-up layer varies with the SHOB and ground surface characteristics, but a few hundred feet are typical for the relatively low yield NTS shots. Techniques for scaling the maximum radius and height of sweep-up layers to larger yields (say  $W \geq 1$  Mt) are still uncertain; however, estimates of ~1000 ft heights for  $W = 1$  Mt seem plausible.

The early-time growth of the sweep-up layer is clearly coupled to the shock wave and air flow characteristics near the ground surface. Figure 8, drawn from photographs at 3 times after burst, illustrates the fact that the sweep-up layer begins forming very quickly behind the shock wave.

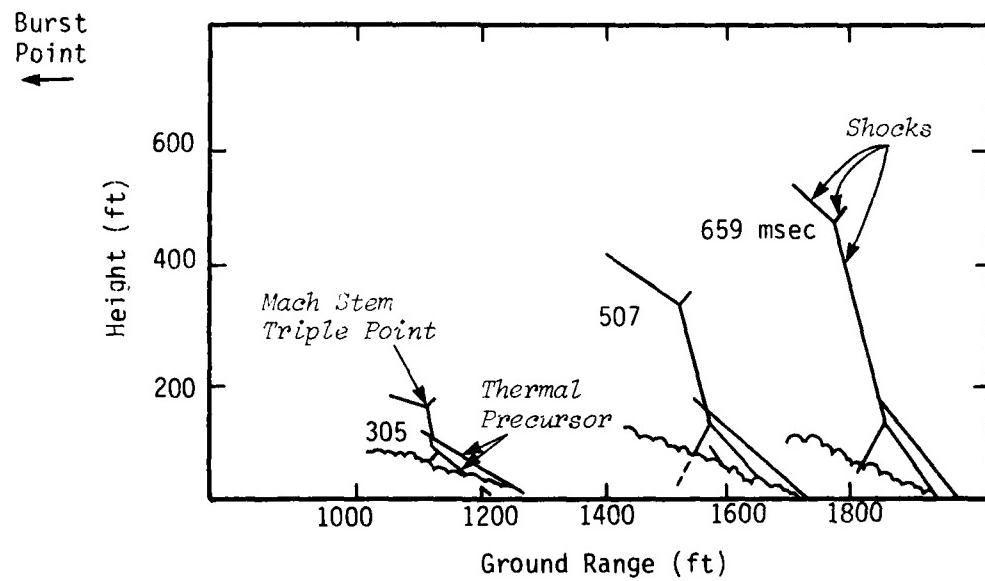


Fig. 8. Tracings from Technical Photography: Shot GRABLE.<sup>7</sup>



Fig. 7. Dusty Sweep-Up Layer for  
Shot GRABLE at  $t \sim 4-5$  Seconds  
( $W = 15$  kt,  $HOB = 160$  m).

The mass concentration or density ( $\rho$ ) of the dust which becomes entrained in the sweep-up layer has been measured in two nuclear tests. For example, a density of about  $2 \times 10^{-3} \text{ gm/cm}^3$  has been measured in shot TEAPOT MET ( $W = 22 \text{ kt}$ , SHOB =  $140 \text{ ft/kt}^{1/3}$ ) at a range corresponding to peak overpressures and particle velocities of  $\sim 20 \text{ psi}$  and  $\sim 600 \text{ ft/sec}$ , respectively. This density was measured at a 3 ft height above the ground surface. Uncertainty factors of about 3 are suspected in this dust density measurement.

The actual concentration profile above the ground surface is not known; however, the following curve shows the general order of magnitude variation expected near the ground surface at a typical range and time in the sweep-up layer.

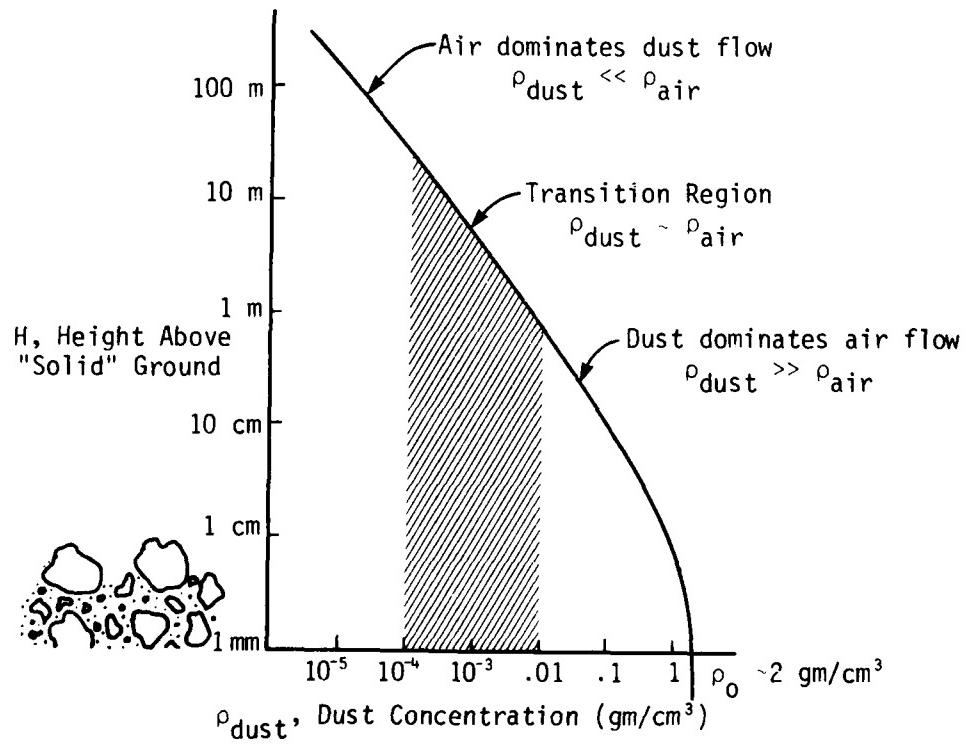
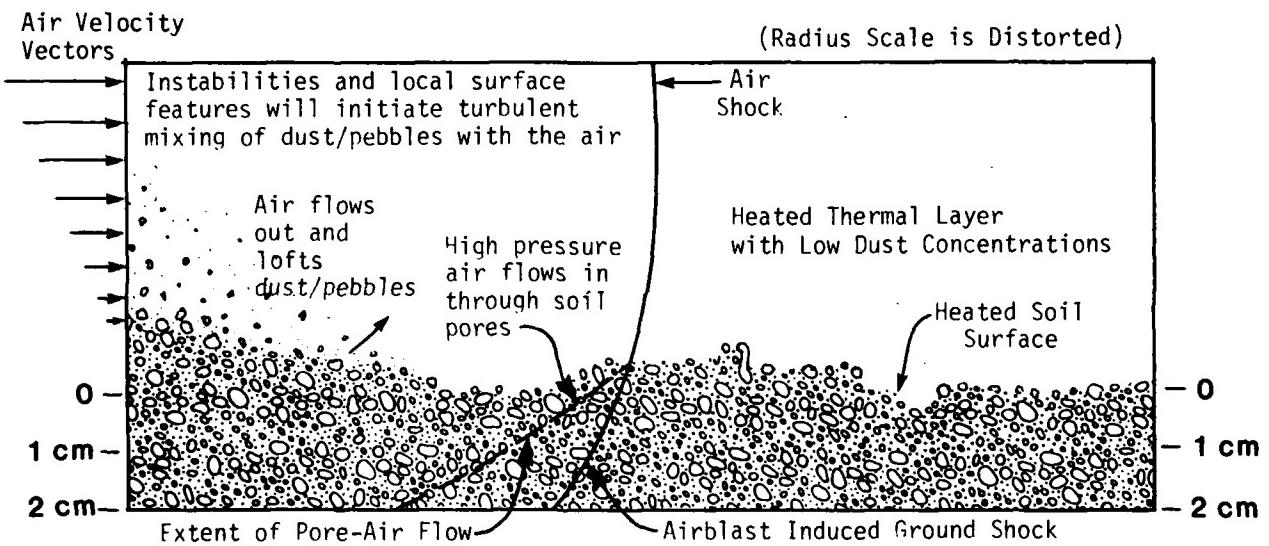


Fig. 9. Estimated Dust Concentration vs Height in the Sweep-Up Layer.

The high dust density region near the original ground surface represents the mass source for the sweep-up layer and ultimately the dust source for the nuclear cloud. The general nature of the interactions between the air and the near-surface material are sketched below:



The multiphase physics (air plus solid/liquid soil particles) in the first 1-10 cm above and below the original ground surface involve viscous stresses which are associated with high gradients in horizontal air velocity near the ground surface. Also, the soil permeability and strength properties will determine the transfer of air momentum and energy to the near-surface soil material.

Immediately behind the shock wave on the ground, there will be some permeation (i.e., diffusion) of air into the porous surface material. The near-surface (~1 cm) permeability variations with depth and range will lead to scouring and lofting of near-surface material on a short-time scale. On a longer time scale, air will permeate through the pores deeper into the soil. As the overpressure on the surface subsequently decays, the air pressure in the pores below the surface will be higher for a time than the pressure being applied onto the surface by the decaying airblast wave. When the air pressures within the pores of the soil exceed the applied overpressure, there is an unbalanced upward force on material near the surface. Thus, soil material is ejected upward into the near-surface air flow and sweep-up layer. This effect has been observed in laboratory experiments and in HE field experiments, and has been theoretically calculated.

The deeper pore-air flow phenomenon may eject ~10 cm of soil material into the sweep-up layer.<sup>8</sup> There is probably a sufficient delay behind the shock wave that this material will not be accelerated to high velocities for a single burst. On the other hand, in a multiple nuclear burst environment, this phenomenon may represent the most significant dust/pebble/rock source for subsequent bursts at some sites. The permeability and cohesive strength of the shocked and heated ground material will probably be the most important site dependent physical characteristic for this phenomenon.

Once the dust/pebbles become lofted above a few meters, the complex air flow field behind the shock wave will dominate the further rise and turbulent diffusion of the dust/pebble mass.

The air flow field behind the near-surface shock wave is quite complicated due to

- unsteady pressure and flow dynamics including Mach stem and thermal precursor formation
- turbulent boundary layer effects involving in-situ bushes, trees, rocks/boulders, and small-scale hills/valleys

Many questions and issues remain unanswered in predicting the formation and nature of the sweep-up layer; Figure 10 summarizes some of these issues.

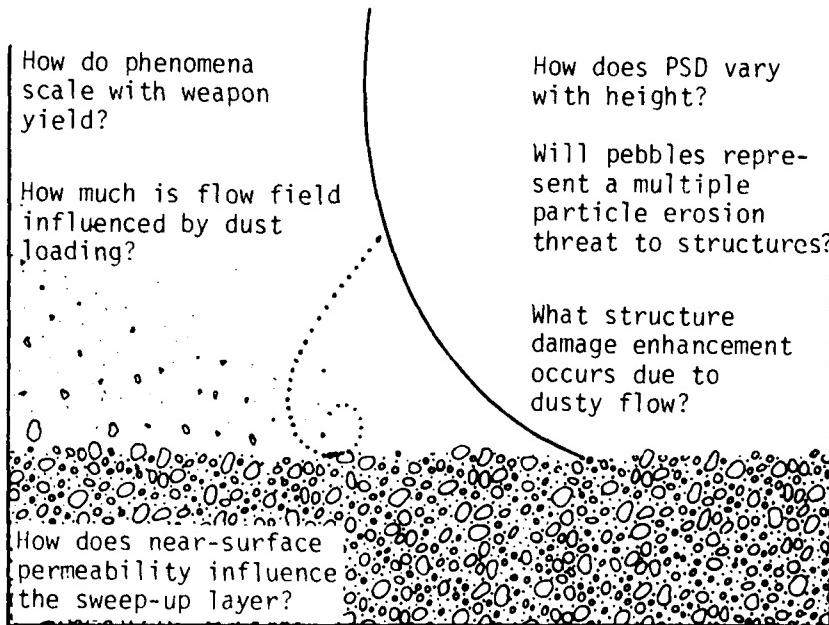


Fig. 10. Issues Concerning Sweep-Up Layer Dust and Pebbles.

### SECTION 3

#### NUCLEAR CLOUD FORMATION

##### 3.1 EARLY-TIME OR FIREBALL PHASE ( $t \leq 10$ seconds for $W = 1$ Mt)

The very early-time phase ( $t \leq 0.1$  second) of a nuclear cloud is characterized by extremely high fireball temperatures ( $10^5$ - $10^6$  °K) and shock wave pressures ( $10^3$ - $10^6$  psi) as indicated in the following Figures 11a and 11b. These curves are from a numerical simulation of a 2 Mt free air burst; this burst is equivalent to a 1 Mt contact surface burst over an idealized rigid surface with no crater and no ejecta.

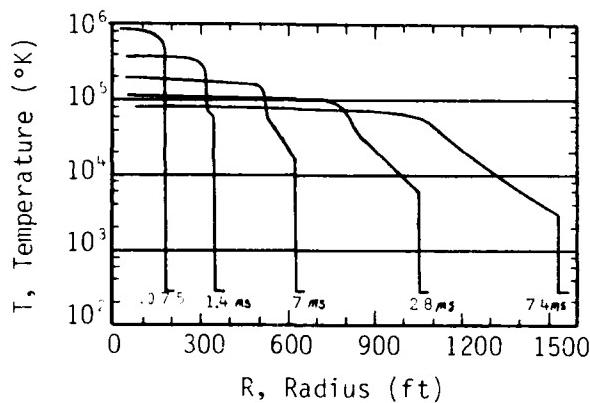


Fig. 11a. Fireball Temperature Versus Radius at Early Times in the Fireball History (1 Mt surface burst).<sup>9</sup>

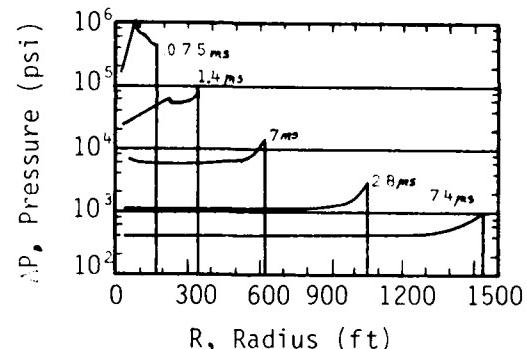


Fig. 11b. Overpressure Versus Radius at Early Times in the Fireball History (1 Mt Surface Burst).<sup>9</sup>

#### EXAMPLE:

At  $t = 28$  milliseconds

$$R_{\text{shock}} = 1,000 \text{ ft}$$

$$T_{\text{shock}} = 8,000^\circ\text{K}$$

$$\Delta P_{\text{shock}} = 4,000 \text{ psi}$$

At the fireball center ( $R = 0$ )

$$T \sim 100,000^\circ\text{K}$$

$$\Delta P \sim 1,000 \text{ psi}$$

The late stages of the fireball growth and associated shock front behavior are indicated on Figures 11c to 11e. The persistence of the high fireball temperatures and low air densities (Figure 11d) set the stage for the buoyancy dominated rise of the hot air and the lofting of entrained dust/debris.

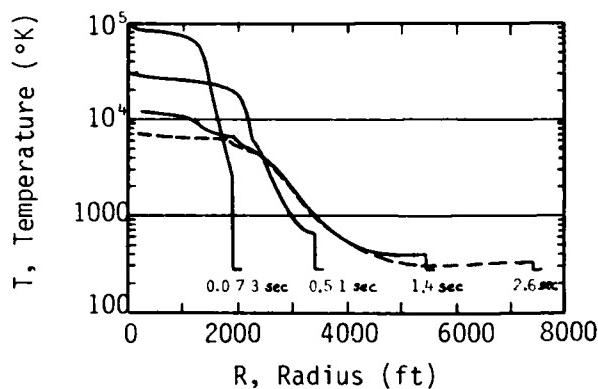


Fig. 11c. Late Fireball Temperature Versus Radius (1 Mt Surface Burst.)

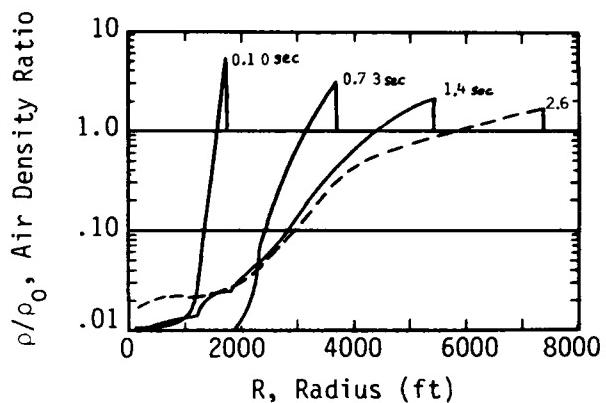


Fig. 11d. Late Fireball Density Ratios ( $\rho_0 \sim 1.2 \times 10^{-3} \text{ gm/cm}^3$ ) Versus Radius (1 Mt Surface Burst.).

Note that the results on Figures 11a to 11e are for an idealized 1-D spherical environment. The following two subsections indicate the effects on the early cloud of crater ejecta for contact surface bursts and the effects of shock wave reflections off the ground surface for HOB detonations.

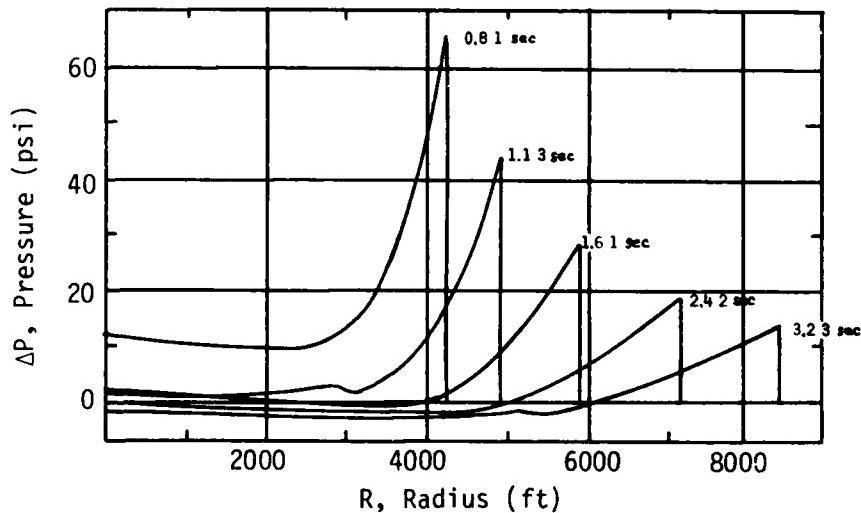


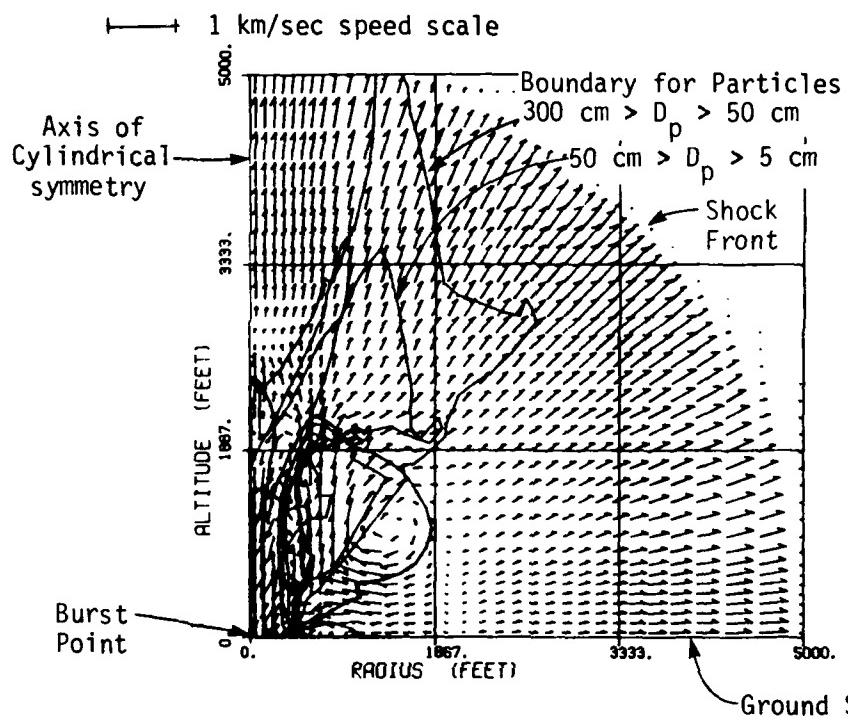
Fig. 11e. Overpressure Versus Radius (1 Mt Surface Burst).<sup>9</sup>

### Early-Time Contact Surface Bursts

The detonation of a nuclear device very near the Earth's surface causes ground material to be thrown from the developing crater into the fireball. The initial pressures in the ground start at about 1000 Mbars ( $\sim 1.5 \times 10^{10}$  psi),<sup>3</sup> and consequently a strong shock wave develops and propagates into the local geologic material as described in Section 2.1. The geology is, therefore, quite important in determining the crater ejecta characteristics.

The ejecta flows into the fireball forming a dynamic multiphase region involving aerodynamic drag and thermal interactions. The thermal interactions include vaporization of solid/liquid material. The vaporization (and later recondensation) of the earth material is potentially significant when comparing shots on land with shots on water.

Figures 12a and 12b show predicted air velocities and temperatures 1 second after a 1 Mt contact surface burst on a soil/rock geology; estimated models were used for crater ejecta characteristics and multiphase drag/thermal interactions. Ejecta mass groups representing different size particles were assumed in the numerical calculations and each of these size groups will interact differently with the air. The boundaries of various size groups are shown on Figure 12a with the larger size particles traveling the larger distances from the burst point. The ejecta influences both the air velocities and temperatures. The temperatures (Figure 12b) are reduced in the interior of the fireball by relatively cold ejecta which mixes with the hotter air.



Velocity vectors show speed (use scale bar at top of figure) and direction of air mass near the base of each vector.

Fig. 12a. Air Velocity Field and Boundaries of Various Particle Size Groups for a 1 Mt Surface Burst at  $t = 1$  sec.

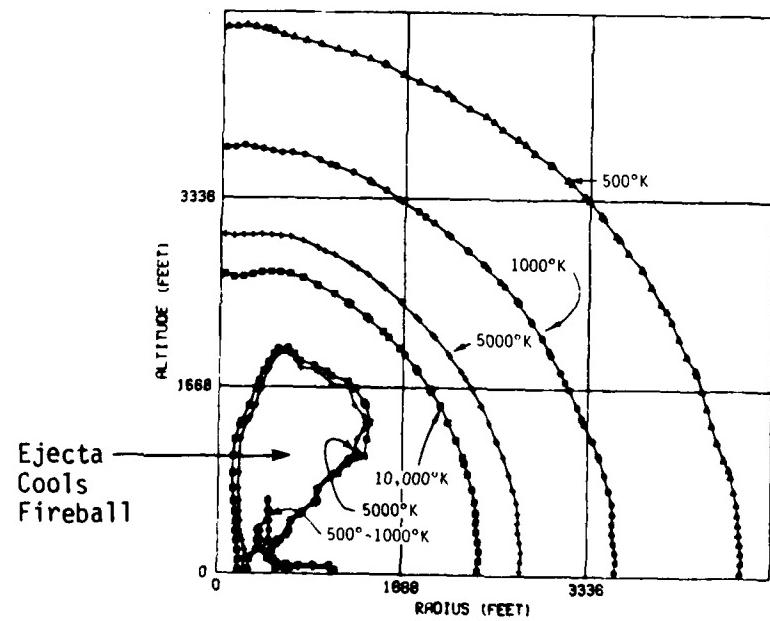


Fig. 12b. Air Temperature Field for a 1 Mt Surface Burst at  $t = 1$  sec.<sup>5</sup>

The predicted dirt concentration field at  $t = 1$  second is shown in Figure 12c. Concentrations of  $10^{-2} \text{ gm/cm}^3$ , which are  $\sim 10$  times denser than ambient air, extend to altitudes of  $\sim 800$  ft in this example. Concentrations of  $10^{-6} \text{ gm/cm}^3$ , which correspond to a relatively high concentration for natural water/ice clouds, extend to  $\sim 2000$  ft in altitude. The mass at heights of  $\sim 3000$  ft consists of the larger diameter particles which are not slowed down by aerodynamic drag.

Figure 12d shows the dirt concentration field at  $t = 14$  seconds. The ejecta plume is near its maximum height of over 10,000 ft. The  $10^{-6} \text{ gm/cm}^3$  contour extends to  $\sim 6000$  ft. Buoyant lofting of the fireball and entrained particles has begun by this time.

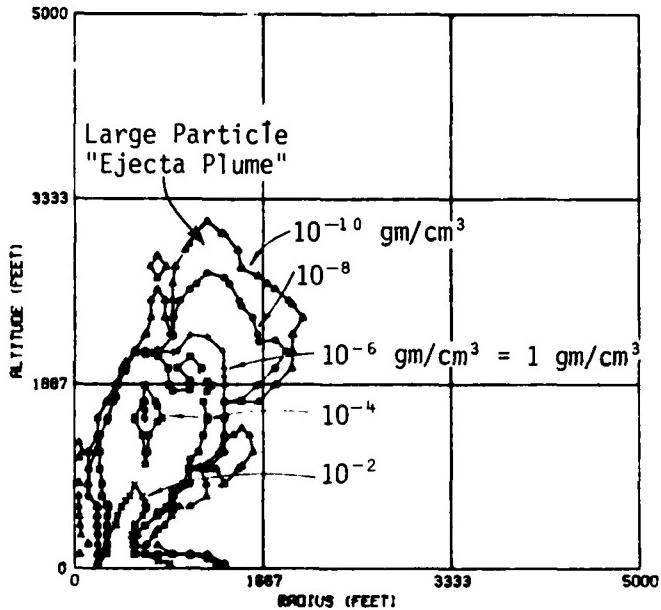


Fig. 12c. Lofted Soil/Rock Concentration Field for a 1 Mt Surface Burst at  $t = 1$  sec.

(Note that radius and altitude scales have been increased by a factor of 3)

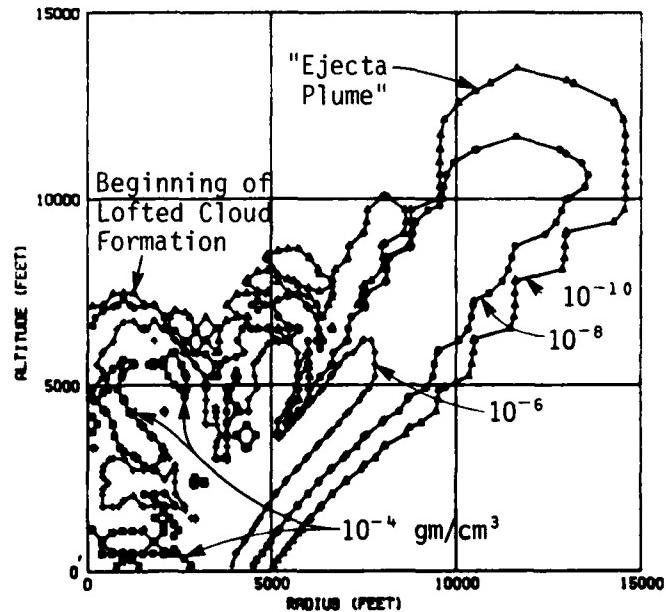


Fig. 12d. Lofted Soil/Rock Concentration Field for a 1 Mt Surface Burst at  $t = 14 \text{ sec.}^5$

### Early-Time HOB Detonations

As the height of the nuclear burst is increased above the Earth's surface, less of the weapon's energy is coupled to ground material and crater ejecta begins to play a smaller and smaller role in forming the nuclear cloud. This transitional region will not be discussed in detail; its characteristics are a combination of the contact surface burst phenomena discussed previously and the non-cratering HOB physical phenomena described in the remainder of this subsection.

When the spherical air shock wave from the burst interacts with the Earth's surface, complex reflected waves are transmitted back into the previously shocked air. These reflected waves strongly influence the subsequent flow field of the air and lofted material. Figures 13a and 13b show an example for shot U/K GRABLE ( $W = 15 \text{ kt}$ , SHOB  $\sim 200 \text{ ft}/\text{kt}^{1/3}$ ) of predicted air velocity vectors just prior to and shortly after shock reflection off the ground.

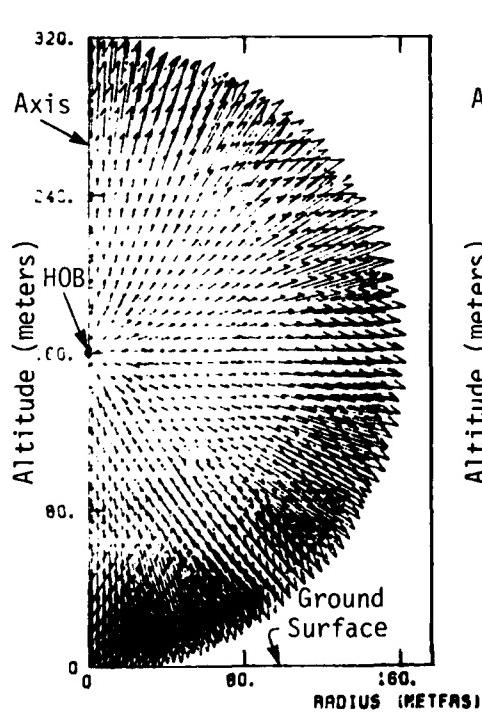


Fig. 13a. Velocity Field at  $t = 0.04$  seconds for U/K GRABLE ( $W = 15 \text{ kt}$ , HOB = 160 m).

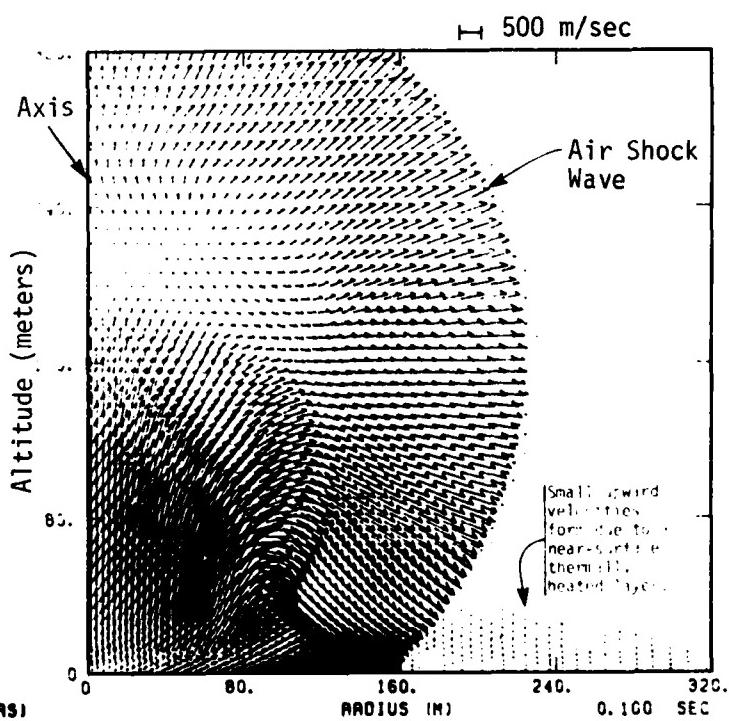


Fig. 13b. Velocity Field at  $t = 0.1$  seconds for U/K GRABLE ( $W = 15 \text{ kt}$ , HOB = 160 m).<sup>10</sup>

Figures 13c and 13d show the air flow fields at  $t = 0.3$  and 0.5 seconds. During this time interval, the primary vortex or torus develops with its center at a radius of about 150 meters and a height of about 240 meters.

Downward directed winds develop above the ground surface and near the axis. A reversal point (RP) in velocity is identified on the adjacent figures. The relatively persistent downward flow will be eventually reversed by buoyancy forces; however, the early entrainment of surface dust/pebbles into the rising fireball will be inhibited.

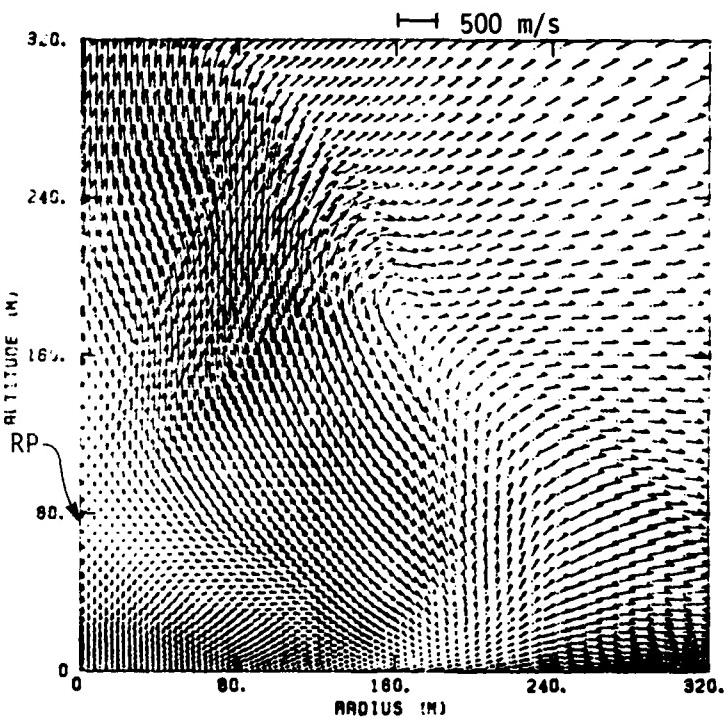


Fig. 13c. Velocity Field at  $t = 0.3$  seconds for U/K GRABLE ( $W = 15$  kt, HOB = 160 m).

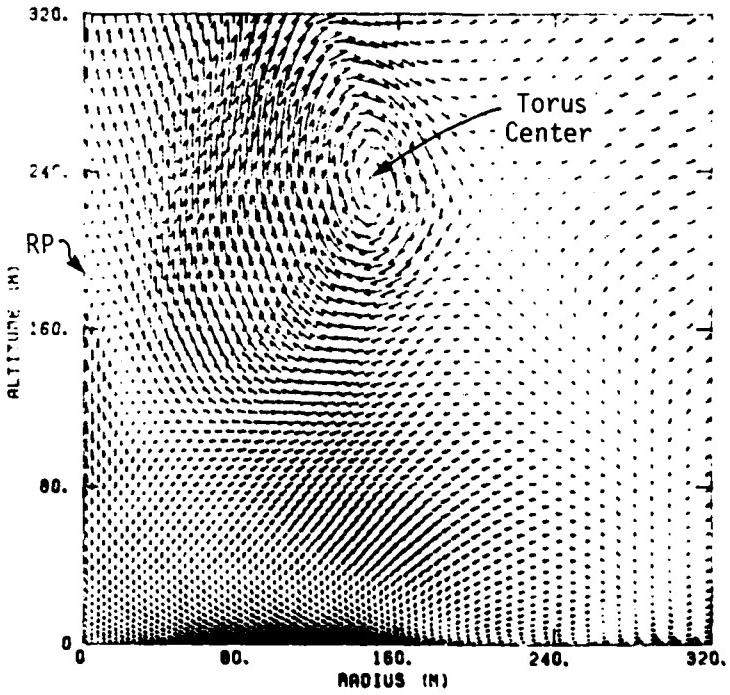


Fig. 13d. Velocity Field at  $t = 0.5$  seconds for U/K GRABLE ( $W = 15$  kt, HOB = 160 m).<sup>10</sup>

Figure 13e shows the predicted air velocity field and the visible fireball shape at  $t = 1.4$  seconds for shot GRABLE. The theoretical/numerical predictions show that the upper fireball region and "tucked in" shape are associated with the torus flow field.

Also shown on Figure 13e is the estimated sweep-up layer height versus radius at this time. Note that the downward velocities below the reversal point (RP) are continuing to suppress the rise of the sweep-up mass near the axis.

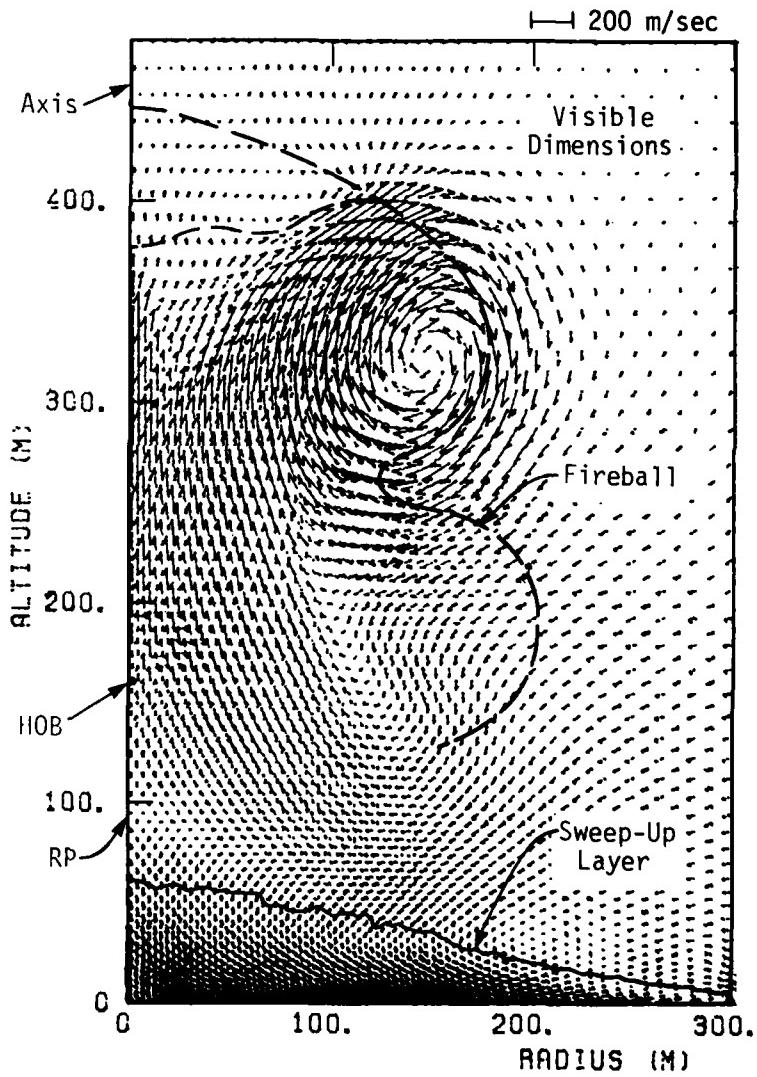


Fig. 13e. Velocity Field at  $t = 1.4$  seconds for U/K GRABLE ( $W = 15$  kt, HOB = 160 m).<sup>10</sup>

### 3.2 CLOUD RISE AND STABILIZATION PHASES (10 sec $\leq$ t $\leq$ 10 min)

After approximately 10-15 seconds for a 1 Mt burst, buoyancy forces begin to dominate the air velocity field characteristics. Buoyancy forces become important because the atmospheric pressure field returns to nearly ambient conditions, but peak air temperatures still exceed 1000°K at 10 seconds. The low density fireball air will accelerate upward causing a vortex air flow field to develop which entrains relatively cool air, earth material, radioactive material, and atmospheric humidity. The vortex flow carries the entrained material through the stem and into the upper portions of the cloud.

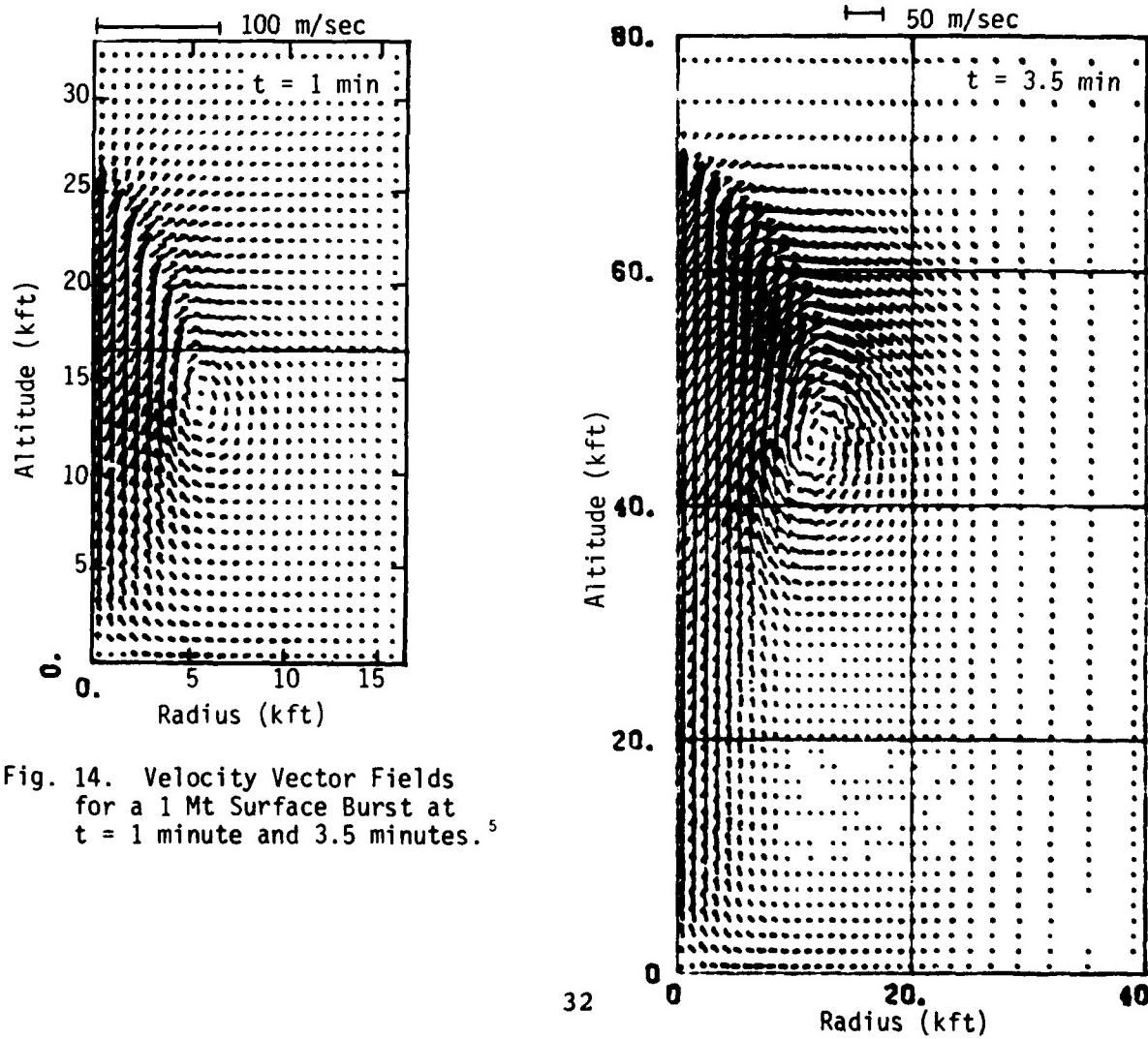


Fig. 14. Velocity Vector Fields  
for a 1 Mt Surface Burst at  
 $t = 1$  minute and 3.5 minutes.<sup>5</sup>

At any time after burst, the amount of earth material and the spatial distribution of this material in the cloud depend on the weapon yield, HOB, and material properties (e.g., size distribution and strength) of the soil/rock. An example follows:

Figure 15 is an estimated spatial distribution of the lofted dust and pebble mass for a 1 Mt contact surface burst at 5 minutes after detonation. Figure 15a shows the mass concentration field. The altitude of the cloud is roughly 70,000 ft with a radius of approximately 20,000 ft at 5 minutes. Note that a vortex region of relatively high density ( $10^{-6} \text{ gm/cm}^3$ ) develops at an altitude of about 50,000 ft and a radius of between 10-15,000 ft.

The nature of the particle size distribution of this mass is presented in Figure 15b, which shows the cumulative lofted mass above any altitude  $Z$  at 5 minutes. Note that most of the mass is in the main cloud (about 210 kilotons of lofted mass above 40,000 ft); the stem contains about 50 kilotons of mass. The cumulative lofted mass is separated into "dust" mass ( $D_p < .5 \text{ cm}$ ) and "pebble" mass ( $.5 < D_p < .5 \text{ cm}$ ). At 5 minutes, over 40% of the lofted mass consists of pebbles. Most of the larger particles have landed on the ground by 5 minutes.

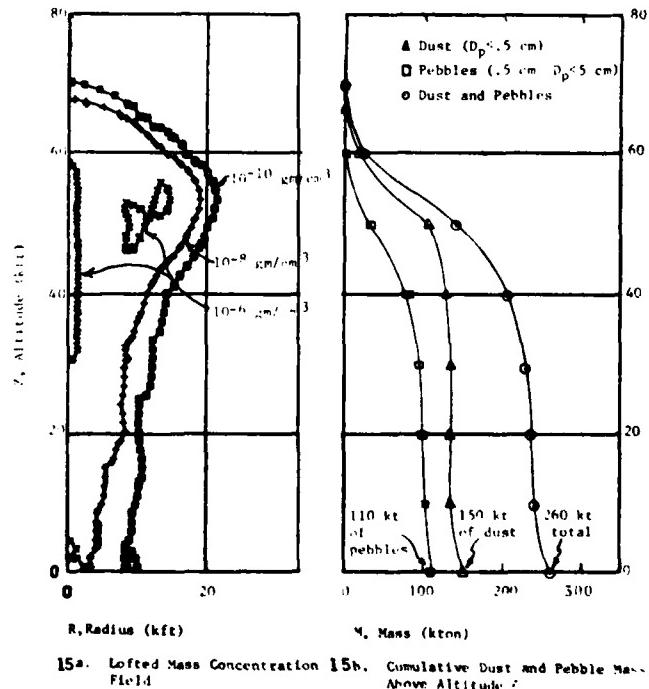


Fig. 15. Lofted Dust/Pebble Spatial (R,Z) Distribution for a 1 Mt Surface Burst at 5 minutes.

The buoyancy dominated phase ends when the cloud reaches its maximum or stabilized cloud altitude. This time is typically 3 to 5 minutes. While the cloud height will stop growing, and in fact decrease in many cases after a few minutes, the cloud diameter will continue to increase; the cloud diameter will increase until about 7-10 minutes even in the absence of atmospheric winds. The increase in cloud diameter is due to the post-stabilization air flow characteristics which occur near the cloud top as illustrated in Figure 16. The reverse vortex forms as a result of downflow of relatively dense air above the cloud top. Relatively dense air from lower altitudes is "pushed" upwards to lower density portions of the atmosphere; when the upward velocities caused by buoyancy get sufficiently small, gravitational forces on the relatively dense air dominate the flow field.

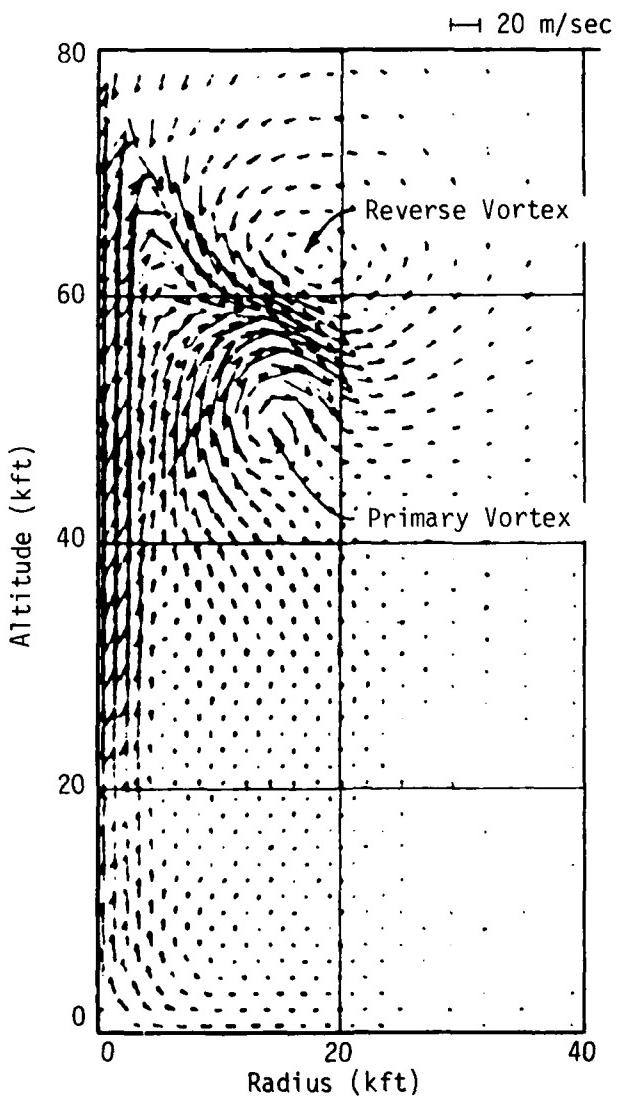


Fig. 16. Air Flow Field at 5 minutes for a 1 Mt Contact Surface Burst.<sup>10</sup>

### 3.3 EXPERIMENTAL CLOUD DIMENSION DATA

Figure 17 is an example of nuclear cloud dimensions versus time for shot CASTLE BRAVO ( $W = 15$  Mt surface burst in the Pacific). At about 5 minutes, the cloud top reaches a maximum altitude of 114 kft; at 9 minutes, the cloud and stem diameter are about 330 kft and 28 kft, respectively.

Figure 18 shows experimental data for cloud altitude versus weapon yield at 1 minute after detonation; Figure 19 shows the maximum (or stabilized) cloud altitude versus yield. The various symbols on these figures indicate whether the shot was conducted in the Pacific or at the Nevada Test Site, and the type of measurement technique used. The scatter in the data is due primarily to variations in atmospheric conditions (e.g., winds, lapse rate, humidity) and/or measurement errors.

Figures 20 and 21 show experimental data for cloud diameter versus weapon yield at 1 minute and 10 minutes after detonation, respectively. Atmospheric winds and turbulence will often cause continued growth in the cloud diameter after 10 minutes.

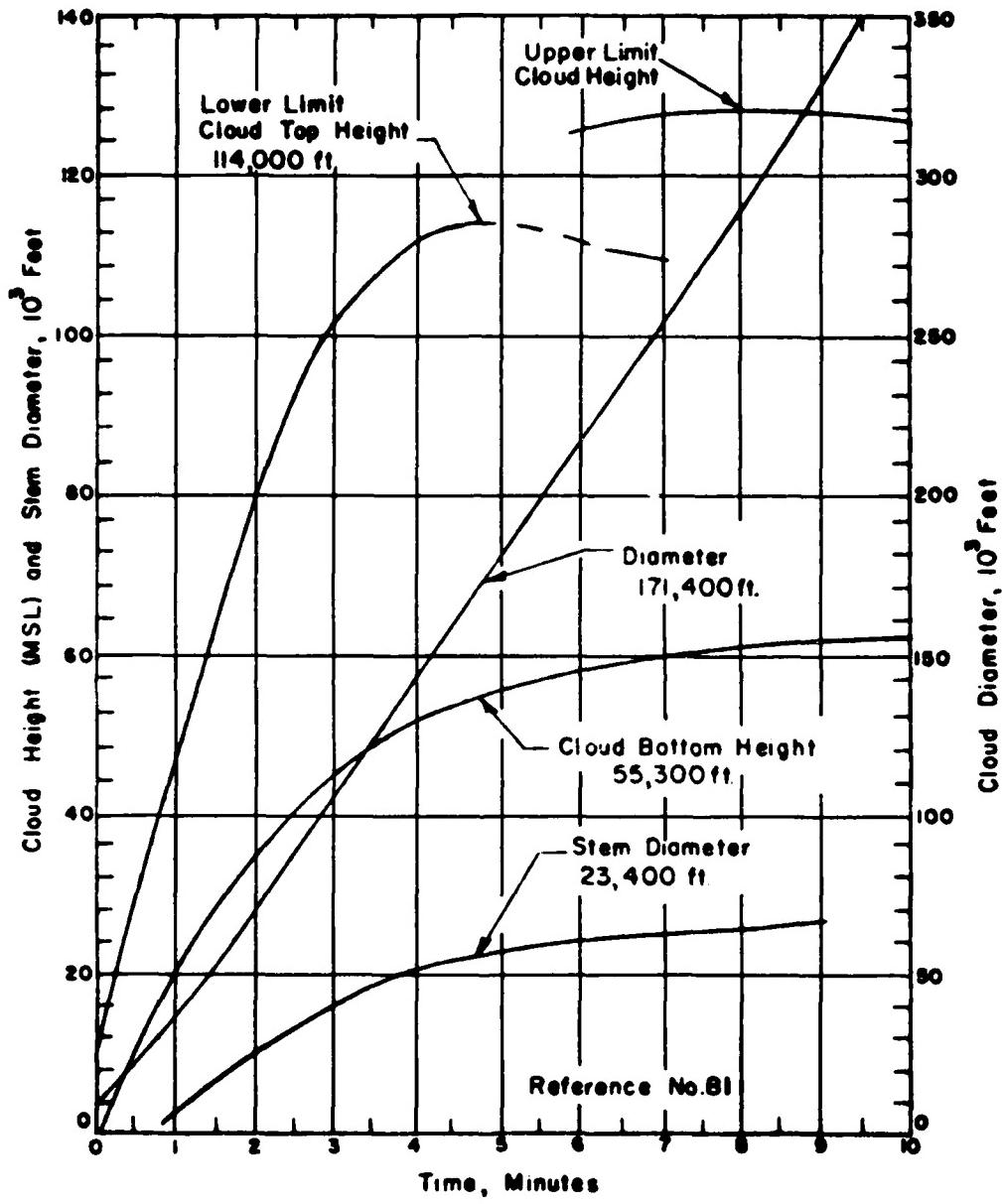


Fig. 17. Cloud Dimensions for Shot CASTLE BRAVO. ( $W = 15$  Mt Surface Burst in the Pacific.)<sup>11</sup>

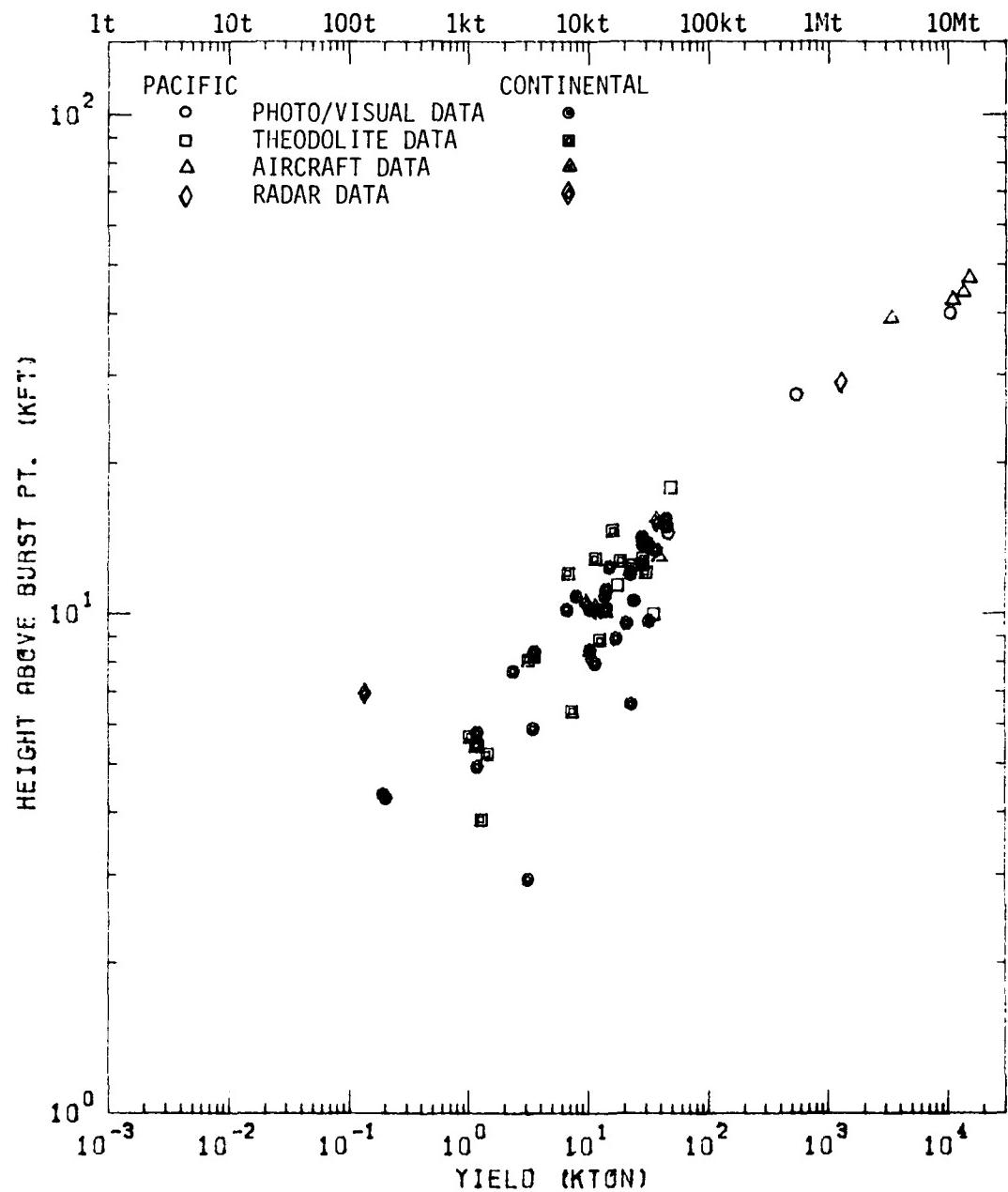


Fig. 18. Cloud Height Above Burst Point Versus Weapon Yield at  $t = 1$  Minute.

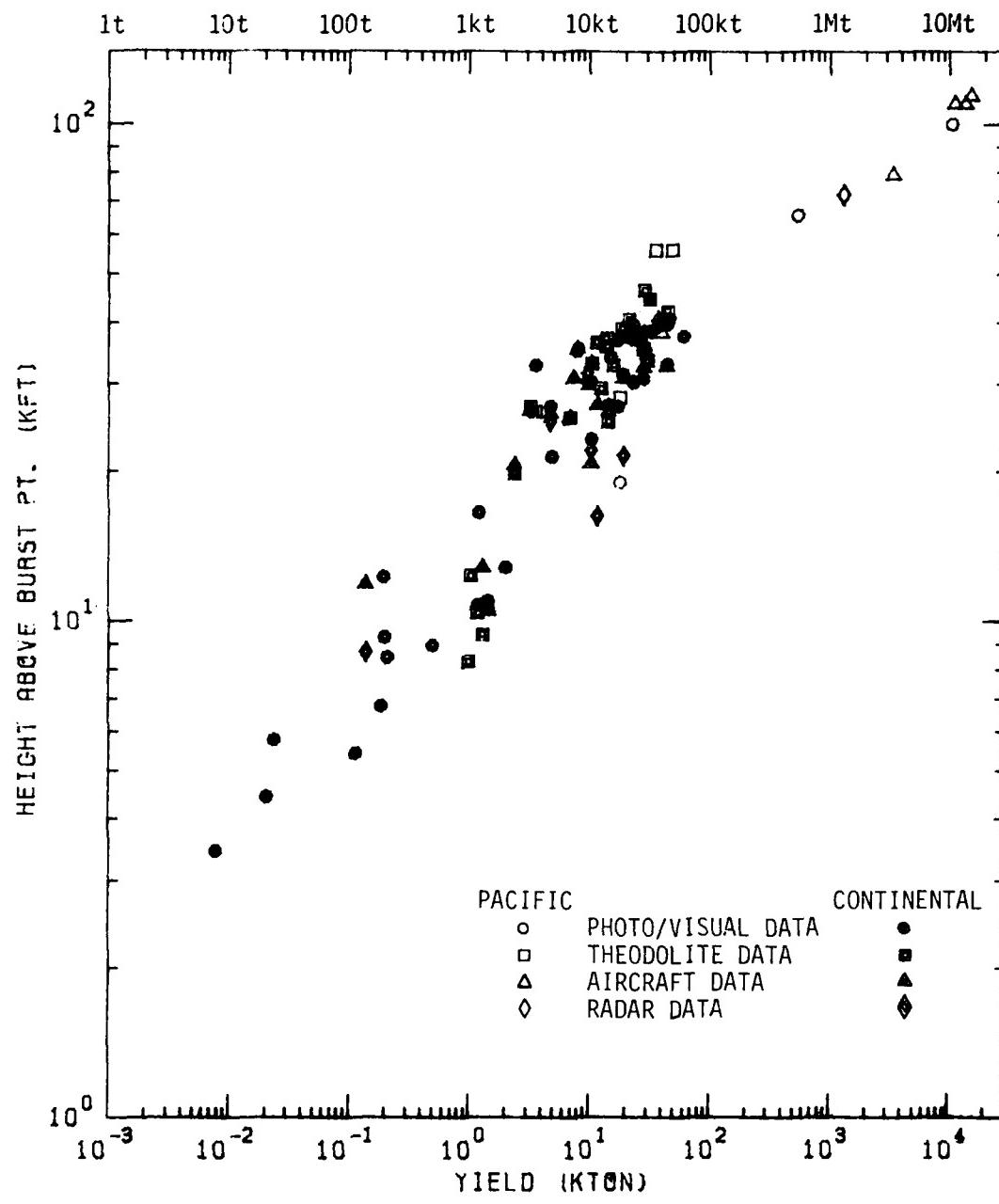


Fig. 19. Cloud Maximum/Stabilization Height Above Burst Point Versus Weapon Yield.<sup>12</sup>

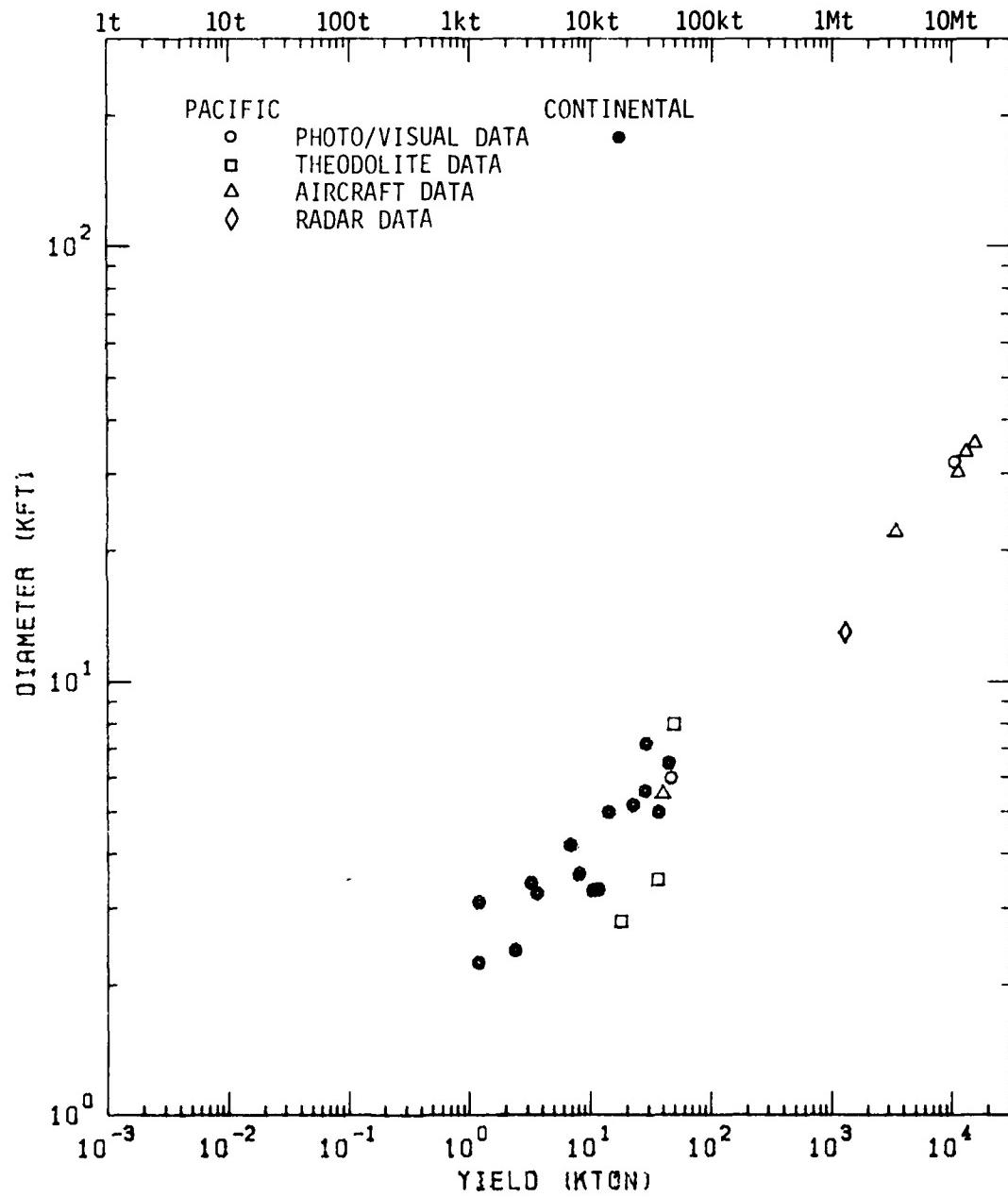


Fig. 20. Cloud Diameter Versus Weapon Yield at  $t = 1$  Minute.<sup>12</sup>

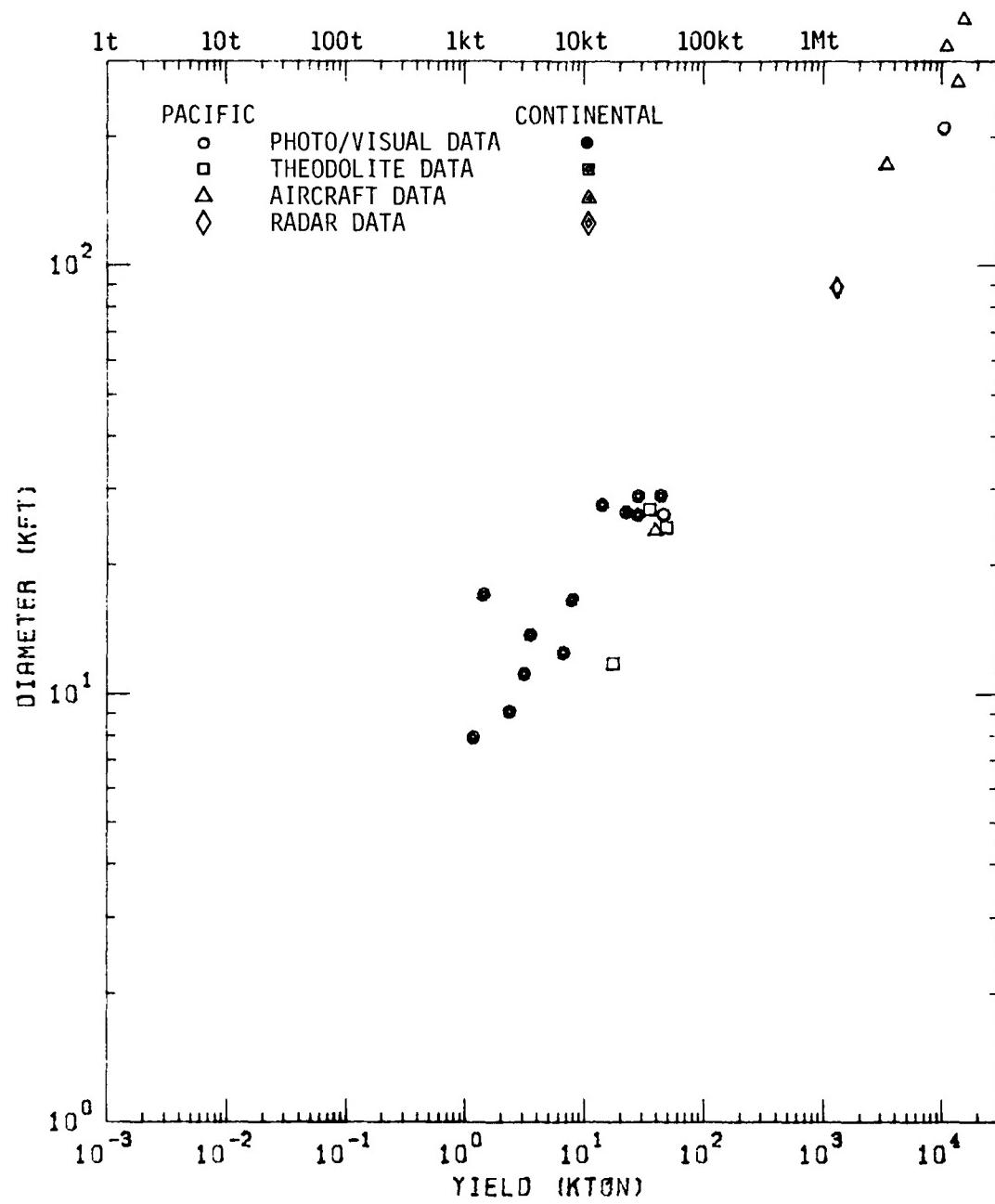


Fig. 21. Cloud Diameter Versus Weapon Yield at  $t = 10$  Minutes.<sup>12</sup>

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